

Thermal properties

1. Determine the dispersion relation:

Write down the equations of motion (masses and springs).

The solutions to these equations will be eigen functions of T

$$\exp\left(i\left(\vec{k} \cdot \vec{a}_1 + \vec{k} \cdot \vec{a}_2 + \vec{k} \cdot \vec{a}_3 - \omega t\right)\right)$$

Substitute the eigen functions of T into the equations of motion to determine the dispersion relation.

2. Determine the density of states numerically from the dispersion relation

$$D(\omega)$$

For every allowed k , find all corresponding values of ω .

long wavelength limit

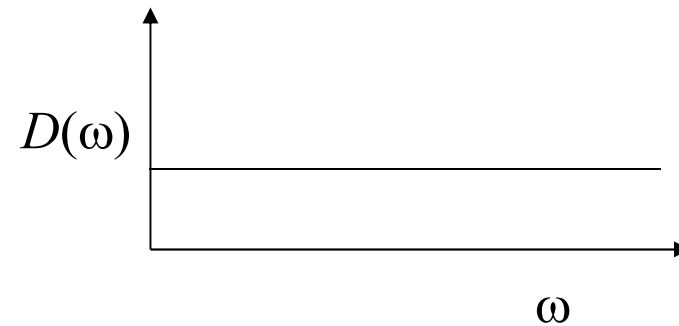
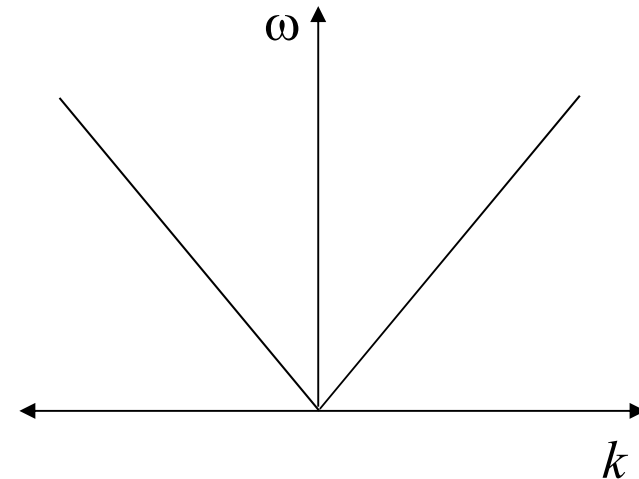
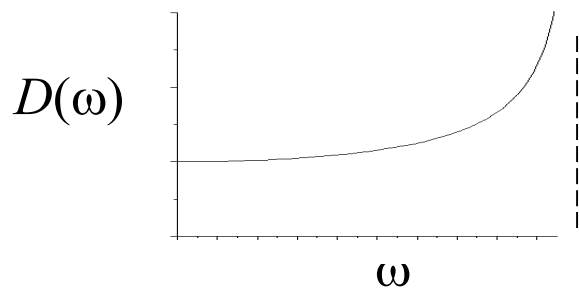
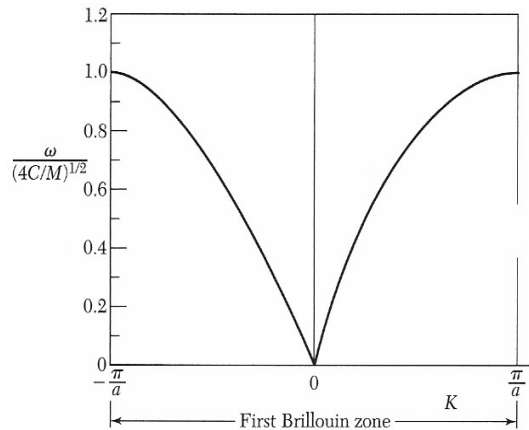
discrete version of wave equation

$$m \frac{d^2 u_s}{dt^2} = C(u_{s+1} - 2u_s + u_{s-1})$$

1-d wave equation

$$\frac{d^2 u}{dt^2} = c^2 \frac{d^2 u}{dx^2}$$

The solutions to the linear chain are the same as the solutions to the wave equation for $|k| \ll \pi/a$.



Phonons - long wavelength, low temperature limit

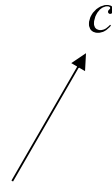
At low T , there are only long wave length states occupied.

3 polarizations

Density of states:
$$D(\omega)d\omega = \frac{3\omega^2}{2c^3\pi^2}d\omega.$$

Specific heat of
insulators at low
temperatures

$$C_v = \frac{24\sigma VT^3}{c}$$



Speed of sound

$$I = \frac{2\pi^5 k_B^4 T^4}{15c^2 h^3} = \sigma T^4 \quad [\text{J m}^{-2} \text{ s}^{-2}]$$

$$u(\lambda) = \frac{8\pi hc}{\lambda^5 \left(\exp\left(\frac{hc}{\lambda k_B T}\right) - 1 \right)} \quad [\text{J/m}^4]$$

$$u = \frac{4\sigma T^4}{c} \quad [\text{J/m}^3]$$

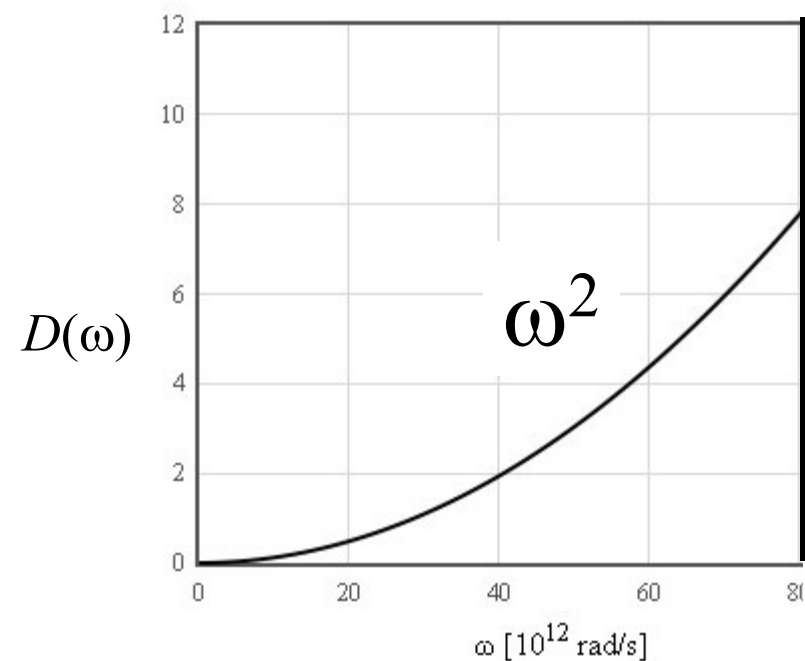
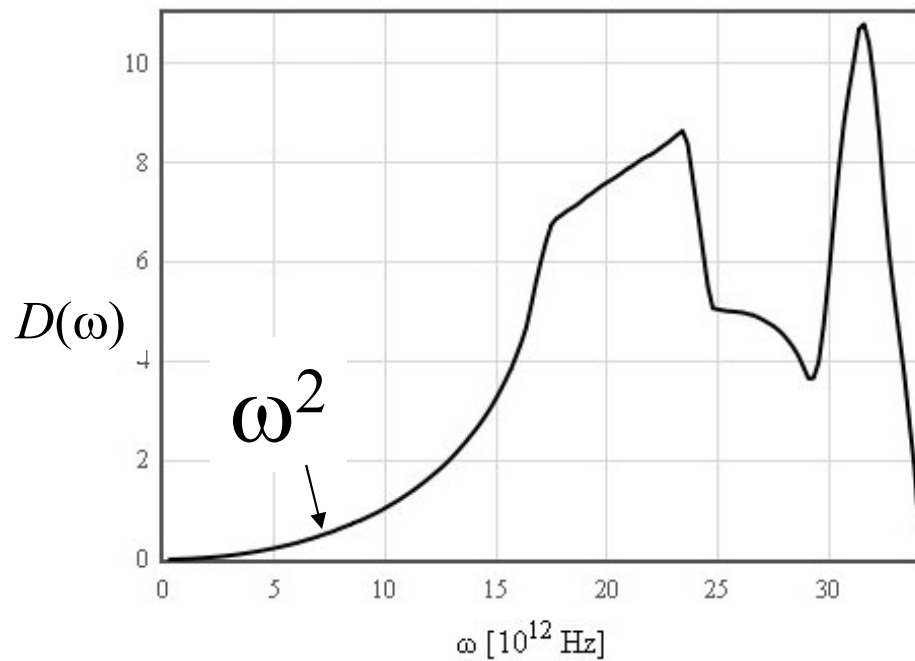
$$c_v = \frac{16\sigma T^3}{c} \quad [\text{J K}^{-1} \text{ m}^{-3}]$$

$$f = \frac{-4\sigma T^4}{3c} \quad [\text{J/m}^3]$$

$$s = \frac{16\sigma T^3}{3c} \quad [\text{J K}^{-1} \text{ m}^{-3}]$$

$$P = \frac{4\sigma T^4}{3c} \quad [\text{N/m}^2]$$

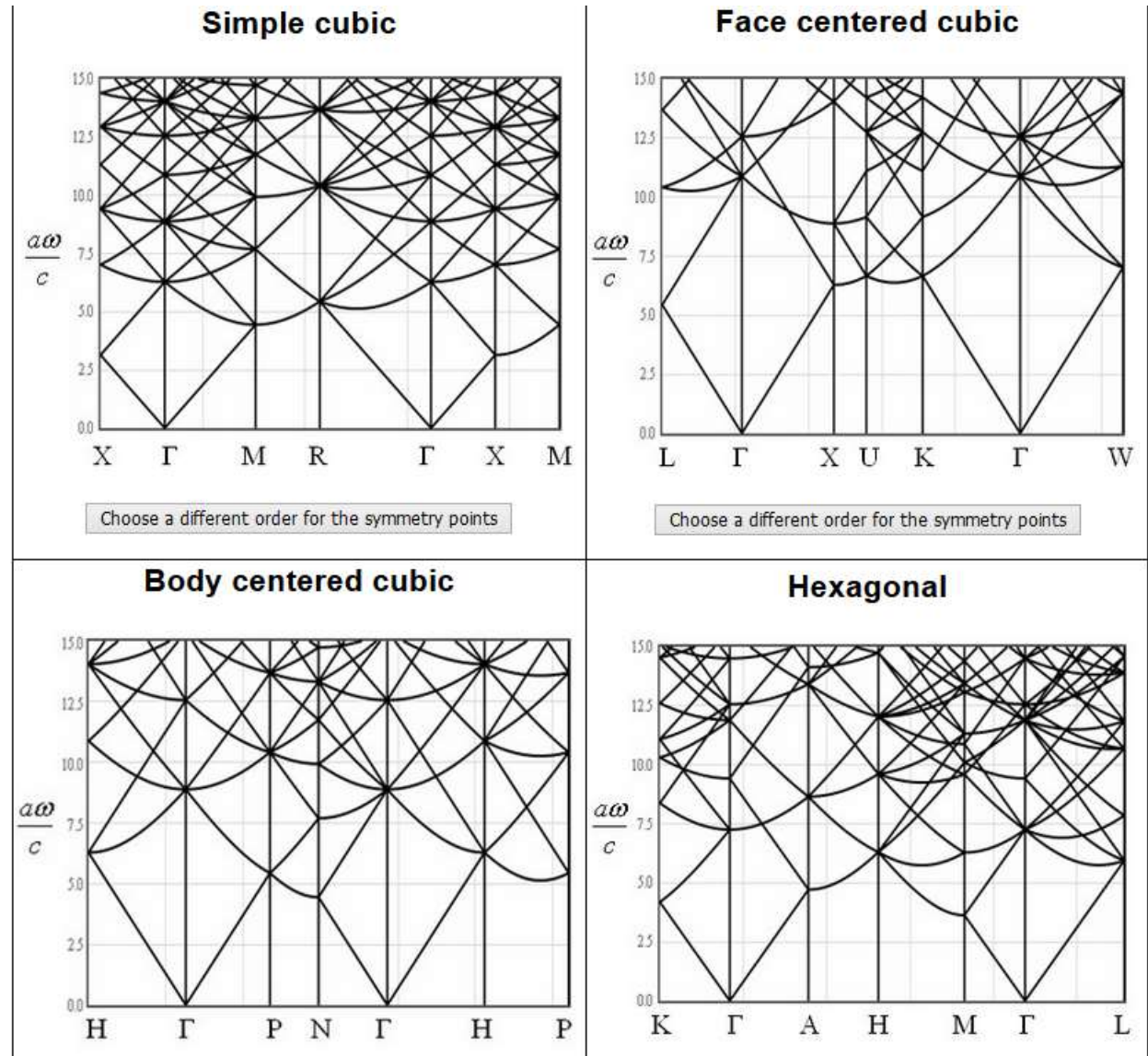
long wavelength, low temperature limit



Empty lattice approximation

Use the speed of sound instead of the speed of light.

3 acoustic branches
3*p* - 3 optical branches



Heat capacity / specific heat

Heat capacity is the measure of the heat energy required to increase the temperature of an object by a certain temperature interval.

Specific heat is the measure of the heat energy required to increase the temperature of a unit quantity of a substance by a certain temperature interval.

For solids, the heat capacity at constant volume and heat capacity at constant pressure are almost the same.

The heat capacity was historically important for understanding solids.

Dulong and Petit (Classical result)

Equipartition: $\frac{1}{2}k_B T$ per quadratic term in energy

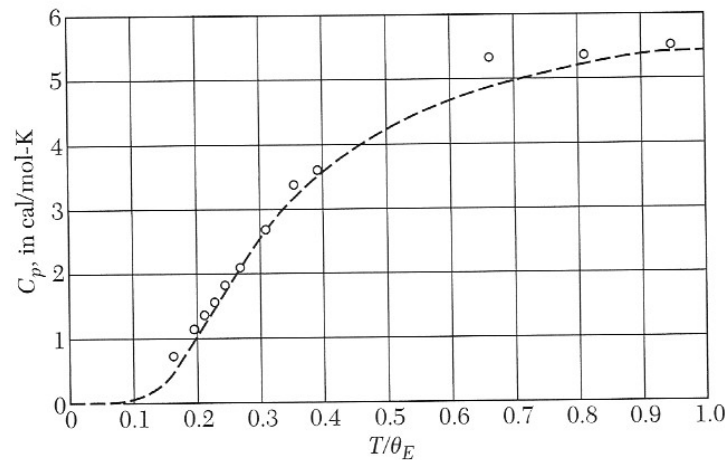
internal energy: $u = 3nk_B T$ N atoms of the crystal

specific heat: $c_v = \frac{du}{dT} = 3nk_B$

experiments: heat capacity goes to zero at zero temperature

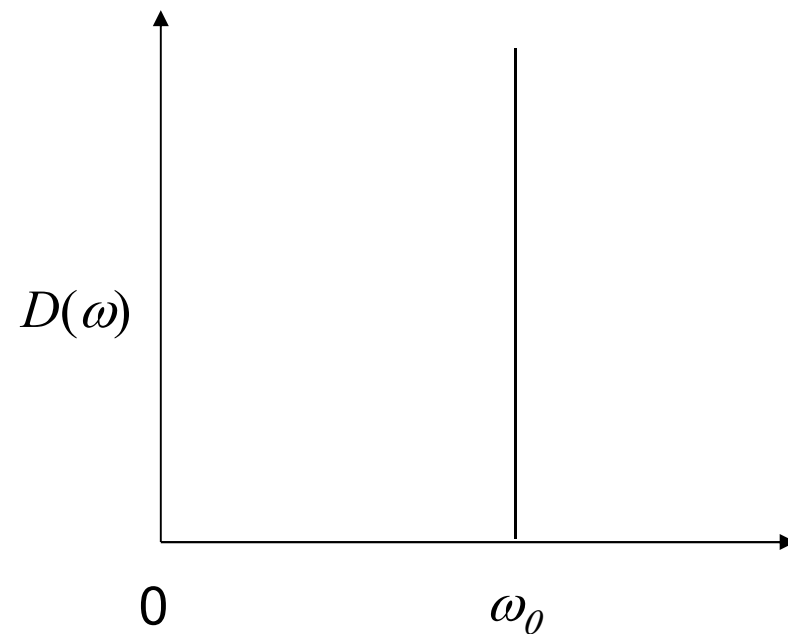
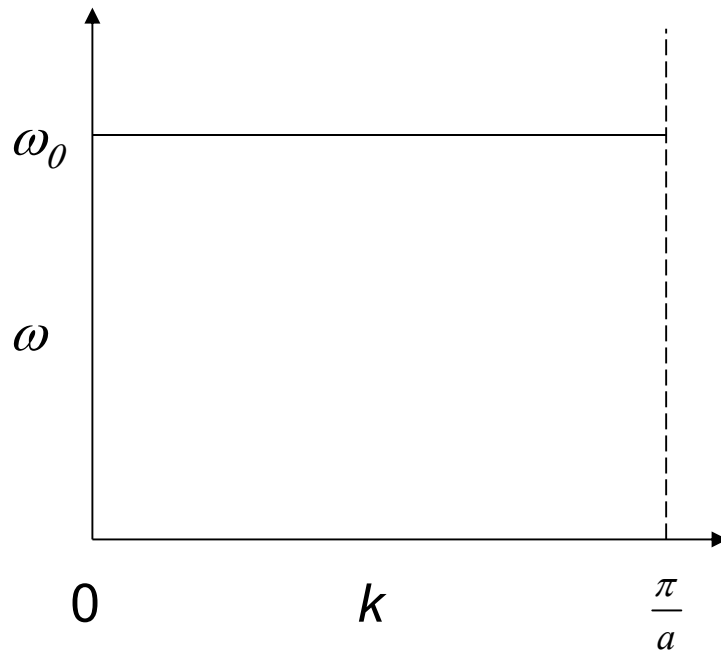


Pierre Louis Dulong



Alexis Therese Petit

Einstein model for specific heat



$$D(\omega) = 3n\delta(\omega - \omega_0)$$

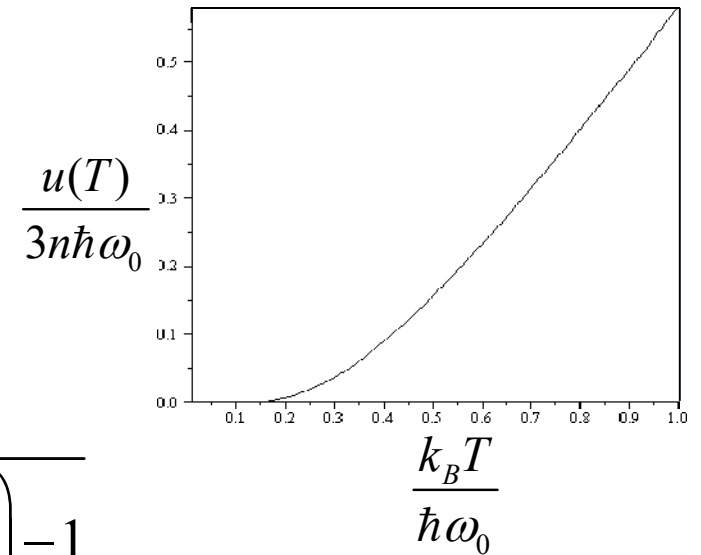
n = density of atoms

$$u(\omega) = D(\omega)\hbar\omega \frac{1}{\exp\left(\frac{\hbar\omega}{k_B T}\right) - 1} = \hbar\omega \frac{3n\delta(\omega - \omega_0)}{\exp\left(\frac{\hbar\omega}{k_B T}\right) - 1}$$

Einstein model for specific heat

$$u(\omega) = \hbar\omega \frac{3n\delta(\omega - \omega_0)}{\exp\left(\frac{\hbar\omega}{k_B T}\right) - 1}$$

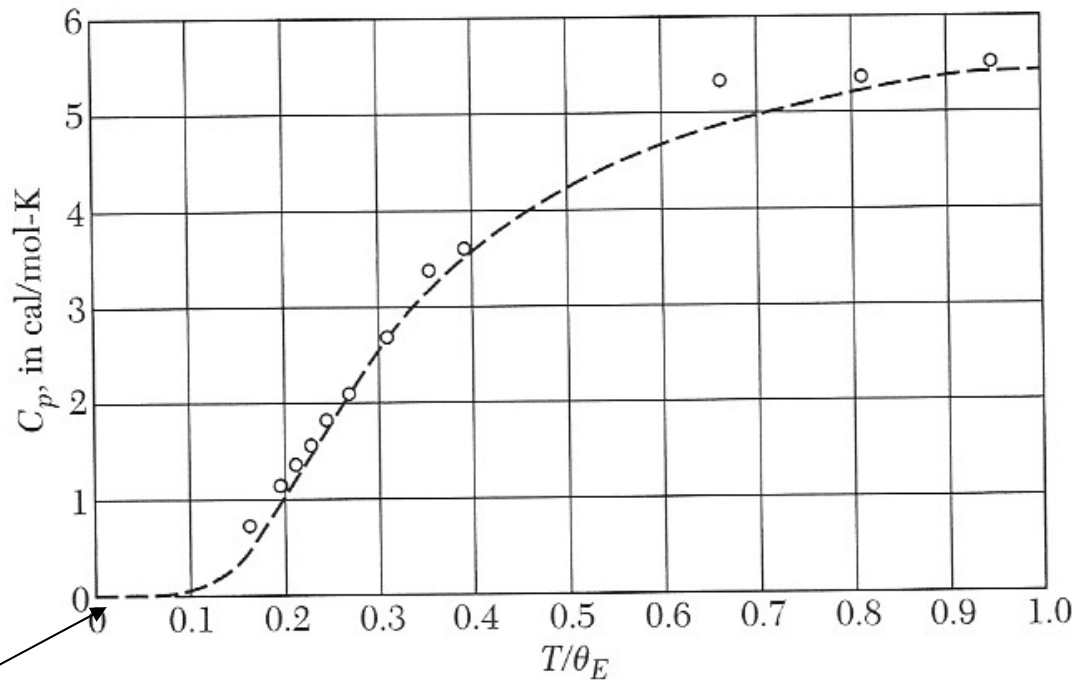
$$u = \int_0^{\infty} u(\omega) d\omega = \int_0^{\infty} 3n\hbar\omega \frac{\delta(\omega - \omega_0) d\omega}{\exp\left(\frac{\hbar\omega}{k_B T}\right) - 1} = \frac{3n\hbar\omega_0}{\exp\left(\frac{\hbar\omega_0}{k_B T}\right) - 1}$$



$$c_v = \frac{du}{dT} = \frac{3n\hbar\omega_0 \frac{\hbar\omega_0}{k_B T^2} \exp\left(\frac{\hbar\omega_0}{k_B T}\right)}{\left(\exp\left(\frac{\hbar\omega_0}{k_B T}\right) - 1\right)^2} = \frac{3nk_B \left(\frac{\hbar\omega_0}{k_B T}\right)^2 \exp\left(\frac{\hbar\omega_0}{k_B T}\right)}{\left(\exp\left(\frac{\hbar\omega_0}{k_B T}\right) - 1\right)^2}$$

Einstein model for specific heat

$$c_v = \frac{3nk_B \left(\frac{\hbar\omega_0}{k_B T} \right)^2 \exp\left(\frac{\hbar\omega_0}{k_B T} \right)}{\left(\exp\left(\frac{\hbar\omega_0}{k_B T} \right) - 1 \right)^2}$$



High temperatures

$$c_v \approx 3nk_B$$

low T does
not fit

$$\theta_E = \frac{\hbar\omega_0}{k_B}$$

Debye model for specific heat



Peter Debye

$$D(\omega) = \frac{3\omega^2}{2\pi^2 c^3}$$

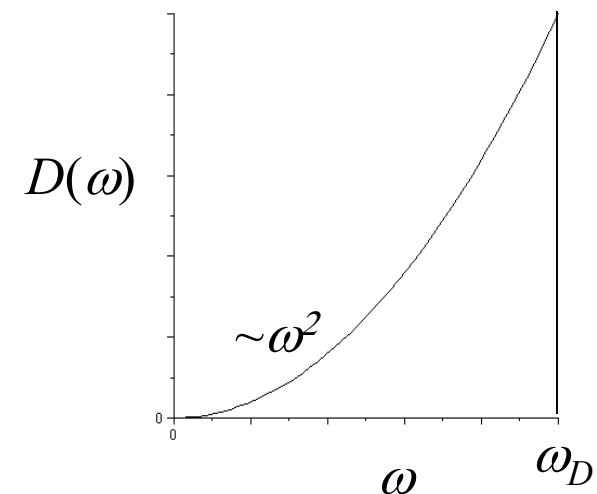
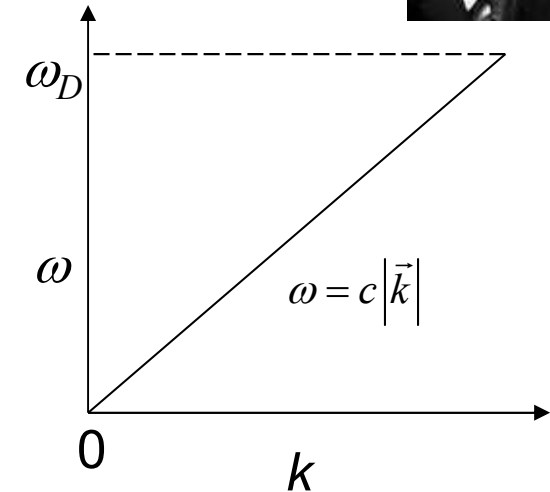
Like blackbody radiation up to a cutoff frequency.

$$V \int_0^{\infty} D(\omega) d\omega = 3N = Na^3 \int_0^{\omega_D} \frac{3\omega^2}{2\pi^2 c^3} d\omega = Na^3 \frac{\omega_D^3}{2\pi^2 c^3}$$

$$\omega_D = \left(\frac{6\pi^2 c^3}{a^3} \right)^{1/3}$$

Debye temperature

$$\hbar\omega_D = k_B \theta_D$$



Debye model for heat capacity

$$u(\omega) = D(\omega)\hbar\omega \frac{1}{\exp\left(\frac{\hbar\omega}{k_B T}\right) - 1} = \frac{3\omega^2}{2\pi^2 c^3} \frac{\hbar\omega}{\exp\left(\frac{\hbar\omega}{k_B T}\right) - 1}$$

$$u = \int_0^{\omega_D} u(\omega) d\omega = \int_0^{\omega_D} \frac{3\omega^2}{2\pi^2 c^3} \frac{\hbar\omega}{\exp\left(\frac{\hbar\omega}{k_B T}\right) - 1} d\omega \approx \int_0^{\infty} \frac{3\omega^2}{2\pi^2 c^3} \frac{\hbar\omega}{\exp\left(\frac{\hbar\omega}{k_B T}\right) - 1} d\omega$$

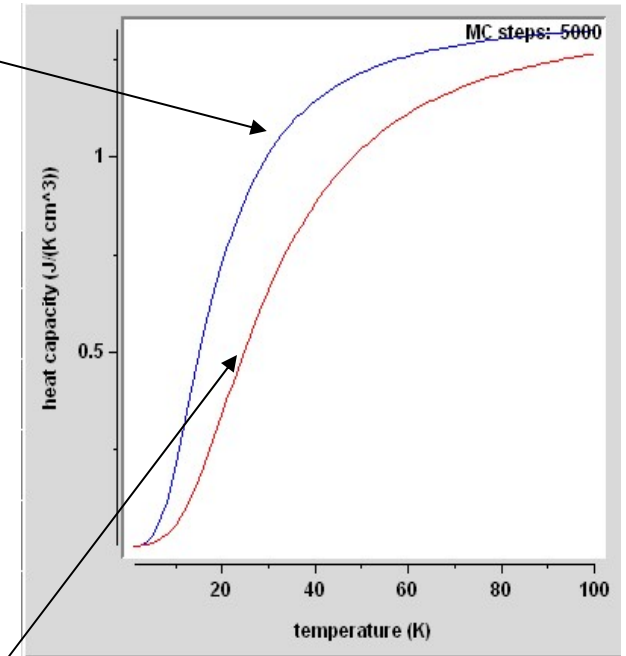
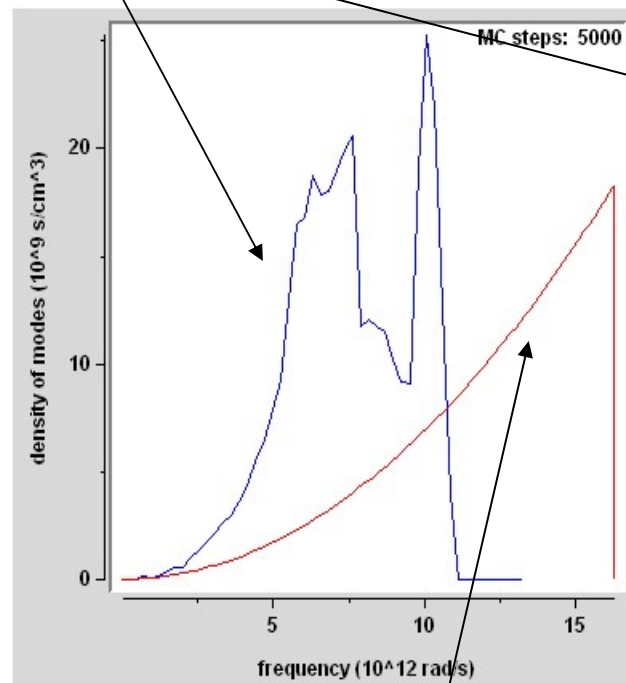
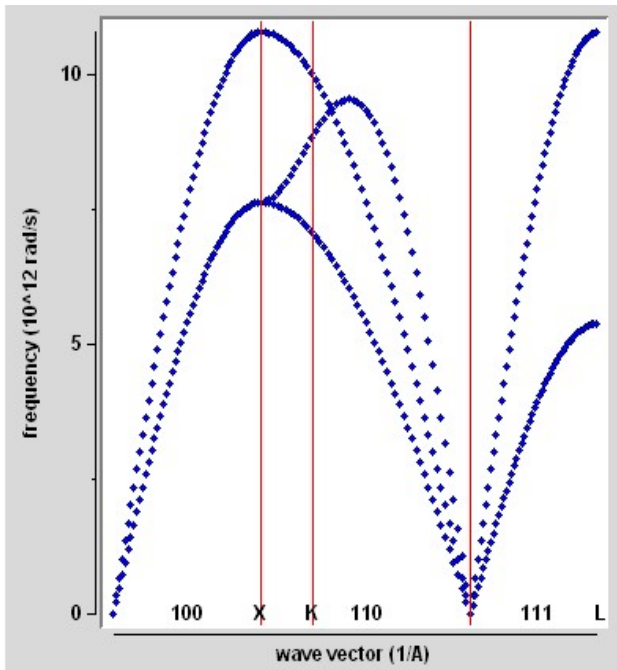
for low T

$$u \approx \frac{3\pi^4}{5} k_B \frac{T^4}{\theta_D^3}$$

$$c_v \approx \frac{12\pi^4}{5} k_B \left(\frac{T}{\theta_D}\right)^3$$

Phonon density of states

fcc phonon density of states



Debye model
(quantized wave
equation with a cut-
off frequency)

Thermal properties

internal energy density $u = \int_0^{\infty} u(\omega) d\omega = \int_0^{\infty} \frac{\hbar\omega D(\omega)}{\exp\left(\frac{\hbar\omega}{k_B T}\right) - 1} d\omega \quad [\text{J/m}^3]$

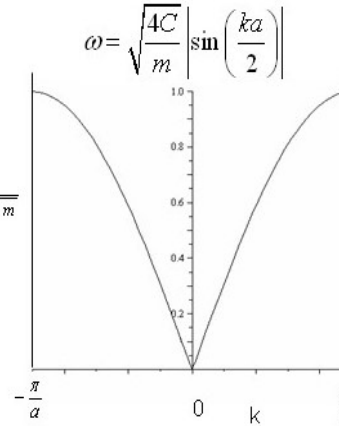
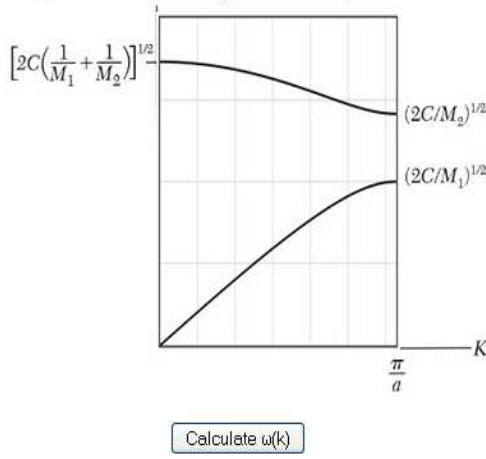
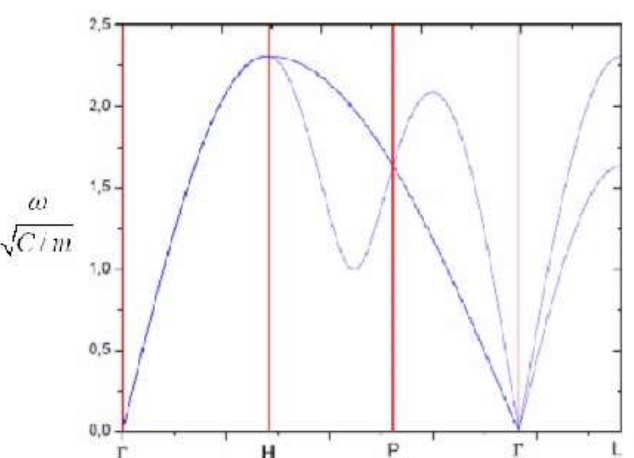
specific heat $c_v = \frac{du}{dT} = \int \left(\frac{\hbar\omega}{T}\right)^2 \frac{D(\omega) \exp\left(\frac{\hbar\omega}{k_B T}\right)}{k_B \left(\exp\left(\frac{\hbar\omega}{k_B T}\right) - 1\right)^2} d\omega \quad [\text{J K}^{-1} \text{ m}^{-3}]$

entropy density $s(T) = \int \frac{c_v}{T} dT = \frac{1}{T} \int_0^{\infty} \frac{\hbar\omega D(\omega)}{\exp\left(\frac{\hbar\omega}{k_B T}\right) - 1} d\omega \quad [\text{J K}^{-1} \text{ m}^{-3}]$

Helmholtz free energy density

$$f(T) = u - Ts = k_B T \int_0^{\infty} D(\omega) \ln \left(1 - \exp\left(\frac{-\hbar\omega}{k_B T}\right) \right) d\omega \quad [\text{J/m}^3]$$

Phonons

	<p style="text-align: center;">Linear Chain</p> $m \frac{d^2 u_s}{dt^2} = C(u_{s+1} - 2u_s + u_{s-1})$	<p style="text-align: center;">Linear chain 2 masses</p> $M_1 \frac{d^2 u_s}{dt^2} = C(v_{s-1} - 2u_s + v_s)$ $M_2 \frac{d^2 v_s}{dt^2} = C(u_s - 2v_s + u_{s+1})$	<p style="text-align: right;"><u>body centered cubic</u></p> $\frac{d^2 u_{lmn}^x}{dt^2} = \frac{C}{\sqrt{3} m} [(u_{l+1m+1n+1}^x - u_{lmn}^x) + (u_{l-1m+1n+1}^x - u_{lmn}^x) + (u_{l+1m-1n+1}^x - u_{lmn}^x) + (u_{l+1m+1n-1}^x - u_{lmn}^x) + (u_{l-1m+1n-1}^x - u_{lmn}^x) + (u_{l+1m+1n+1}^y - u_{lmn}^y) - (u_{l-1m+1n+1}^y - u_{lmn}^y) - (u_{l+1m-1n+1}^y - u_{lmn}^y) - (u_{l+1m+1n-1}^y - u_{lmn}^y) + (u_{l-1m+1n-1}^y - u_{lmn}^y) + (u_{l+1m-1n-1}^y - u_{lmn}^y) + (u_{l+1m+1n+1}^z - u_{lmn}^z) - (u_{l-1m+1n+1}^z - u_{lmn}^z) - (u_{l+1m-1n+1}^z - u_{lmn}^z) - (u_{l+1m+1n-1}^z - u_{lmn}^z) + (u_{l-1m+1n-1}^z - u_{lmn}^z) - (u_{l+1m-1n-1}^z - u_{lmn}^z)]$ <p style="text-align: right;">And similar expressions for the y and z</p>
<p>Eigenfunction solutions</p>	$u_s = A_x e^{i(ksa - \omega t)}$	$u_s = u e^{i(ksa - \omega t)}$ $v_s = v e^{i(ksa - \omega t)}$	$u_{lmn}^x = u \frac{x}{k} e^{i(l \vec{k} \cdot \vec{a}_1 + m \vec{k} \cdot \vec{a}_2 + n \vec{k} \cdot \vec{a}_3)} = u \frac{x}{k} e^{i(-l - m - n) \frac{\pi}{a}}$ <p style="text-align: right;">And similar expressions for the y and z</p>
<p>Dispersion relation</p>	$\omega = \sqrt{\frac{4C}{m}} \left \sin\left(\frac{ka}{2}\right) \right $ 	$\omega^2 = C \left(\frac{1}{M_1} + \frac{1}{M_2} \right) \pm C \sqrt{\left(\frac{1}{M_1} + \frac{1}{M_2} \right)^2 - \frac{4 \sin^2\left(\frac{ka}{2}\right)}{M_1 M_2}}$ 	<p style="text-align: right;">The dispersive</p> $\begin{aligned} & 4 - \cos\left(\frac{\alpha}{2}(k_x + k_y + k_z)\right) - \cos\left(\frac{\alpha}{2}(3k_x - k_y - k_z)\right) && -\cos\left(\frac{\alpha}{2}(k_x + k_y - k_z)\right) \\ & -\cos\left(\frac{\alpha}{2}(-k_x + 3k_y - k_z)\right) - \cos\left(\frac{\alpha}{2}(-k_x - k_y + 3k_z)\right) - \frac{m\omega^2}{\sqrt{3}C} && +\cos\left(\frac{\alpha}{2}(-k_x + 3k_y - k_z)\right) \\ & -\cos\left(\frac{\alpha}{2}(k_x + k_y + k_z)\right) + \cos\left(\frac{\alpha}{2}(3k_x - k_y - k_z)\right) && 4 - \cos\left(\frac{\alpha}{2}(k_x + k_y - k_z)\right) \\ & +\cos\left(\frac{\alpha}{2}(-k_x + 3k_y - k_z)\right) - \cos\left(\frac{\alpha}{2}(-k_x - k_y + 3k_z)\right) && -\cos\left(\frac{\alpha}{2}(-k_x + 3k_y - k_z)\right) \\ & -\cos\left(\frac{\alpha}{2}(k_x + k_y + k_z)\right) + \cos\left(\frac{\alpha}{2}(3k_x - k_y - k_z)\right) && -\cos\left(\frac{\alpha}{2}(k_x + k_y - k_z)\right) \\ & -\cos\left(\frac{\alpha}{2}(-k_x + 3k_y - k_z)\right) + \cos\left(\frac{\alpha}{2}(-k_x - k_y + 3k_z)\right) && +\cos\left(\frac{\alpha}{2}(-k_x + 3k_y - k_z)\right) \end{aligned}$ 
<p>Density of states $D(k)$</p>	$D(k) = \frac{1}{\pi}$	$D(k) = \frac{1}{\pi}$	$D(k) = \frac{3k^2}{2\pi^2}$

Quartz

α -Quartz
trigonal
2.65 g/cm³

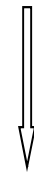
573°C
⇒

β -Quartz
hexagonal
2.53 g/cm³

870°C
⇒

β -Tridymite
hexagonal
2.25 g/cm³

1470°C



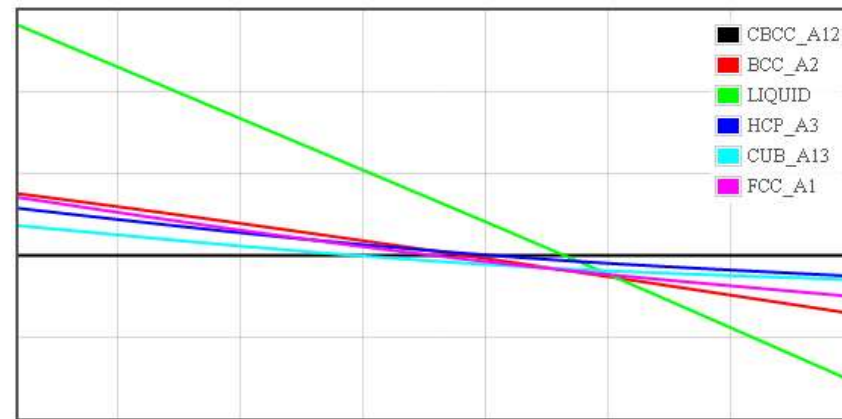
β -Cristobalite
cubic
2.20 g/cm³

Silica Melt

1705°C
⇐



f



T

Phonon student projects

Calculate a dispersion relation including next nearest neighbors.

Write a javascript program that plots the phonon dispersion relation in an arbitrary direction.

Calculate one column of the phonon table: hcp, NaCl, CsCl, ZnS, diamond, ...

Calculate the temperatures at which ZnO goes through a phase transition.

Waves and particles

The eigen function solutions of the wave equation are plane waves. The scattering time is one over the rate for scattering from a given plane wave solution to any other.

Phonons are particles. The scattering time is the time before the phonons scatter and randomly change energy and momentum.

$$\vec{p} = \hbar \vec{k}$$

The average time between scattering events is $\tau_{sc} = 1/\Gamma$

Phonon scattering

Scattering randomizes the momentum of the phonons.

$$H = H_{HO} + H_1$$

Transition rates determined by Fermi's golden rule

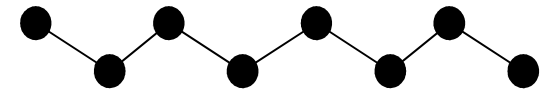
$$\Gamma_{i \rightarrow f} = \frac{2\pi}{\hbar} \left| \langle \psi_f | H_1 | \psi_i \rangle \right|^2 \delta(E_f - E_i)$$

Any process (3 phonon, 4 phonon, 5 phonon. ...) that conserves energy and momentum is allowed.

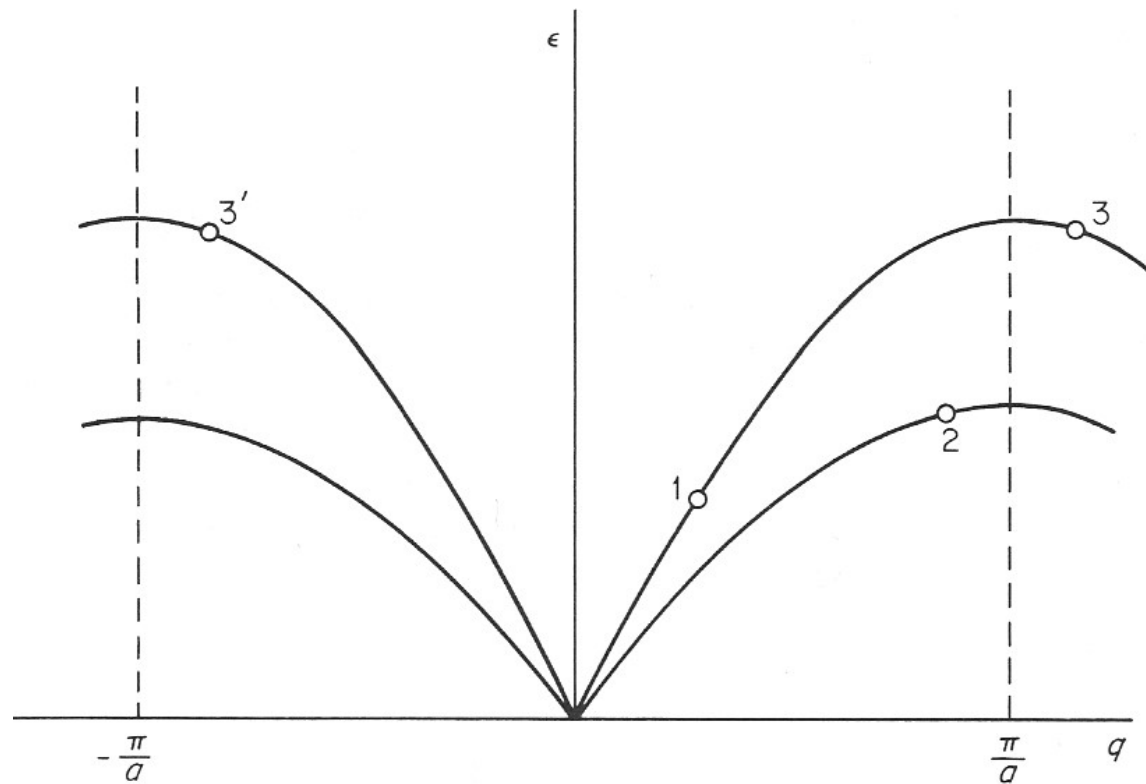
Results in attenuation of acoustic waves

Umklapp Processes

Three phonon scattering



$$\hbar\vec{k}_1 + \hbar\vec{k}_2 = \hbar\vec{k}_3 + \hbar\vec{G}$$

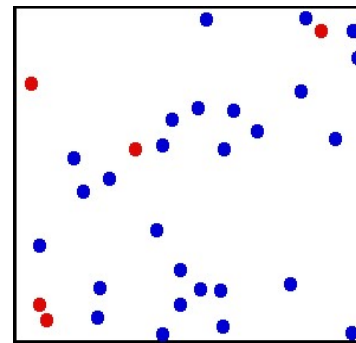
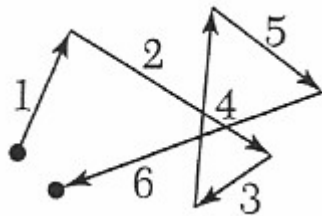


from: Hall, Solid State Physics

Heat transport (Kinetic theory)

Treat phonons as an ideal gas of particles that are confined to the volume of the solid.

Phonons move at the speed of sound. They scatter due to imperfections in the lattice and anharmonic terms in the Hamiltonian.



The average time between scattering events is τ_{sc}

The average distance traveled between scattering events is the mean free path: $l = v\tau_{sc} \sim 10 \text{ nm}$

Diffusion equation/ heat equation

Diffusion constant $\frac{dn}{dt} = -D\nabla^2 n$

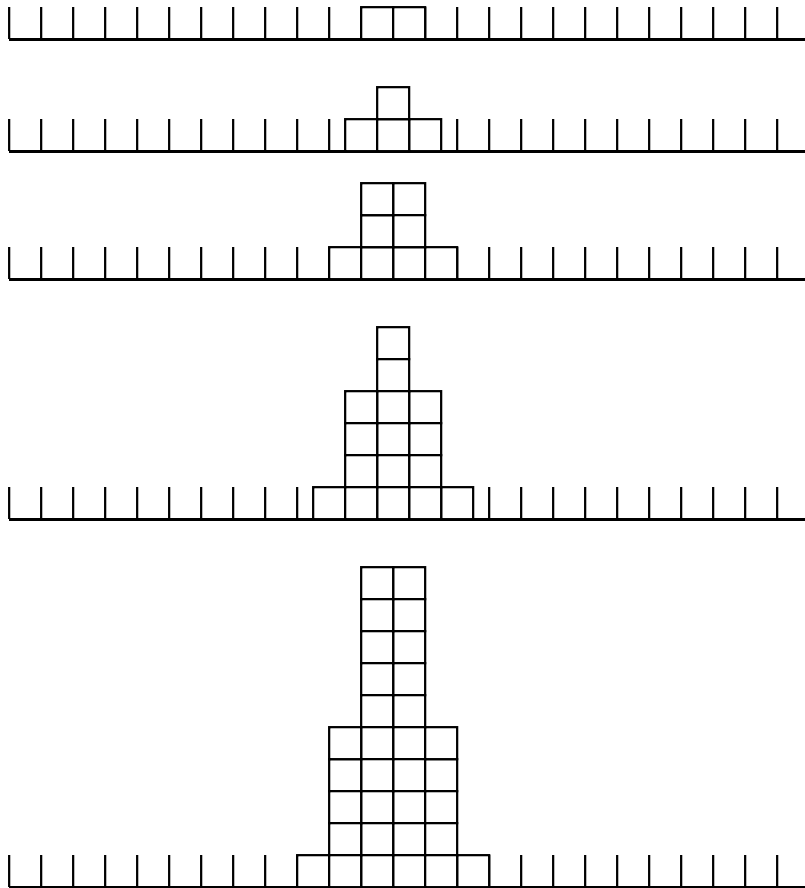
Fick's law $\vec{j} = -D\nabla n$

Continuity equation $\frac{dn}{dt} = \nabla \cdot \vec{j}$



$$n = \frac{1}{\sqrt{4\pi Dt}} \exp\left(\frac{-r^2}{4Dt}\right)$$

Random walk



$$\frac{\Delta n_s}{\Delta t} = n_{s+1} - 2n_s + n_{s-1}$$



$$\exp(-x^2)$$

Central limit theorem: A function convolved with itself many times forms a Gaussian

Thermal conductivity

$$\vec{j}_U = \bar{E} \vec{j}$$

Average particle energy

$$u = \bar{E} n$$

internal energy density

$$\vec{j}_U = -\bar{E} D \nabla n = -D \nabla u$$

$$\vec{j}_U = -D \frac{du}{dT} \nabla T = -D c_v \nabla T$$

$$\vec{j}_U = -K \nabla T$$

Thermal conductivity

$$K = D c_v$$

$$K \rightarrow 0 \quad \text{as} \quad T \rightarrow 0$$